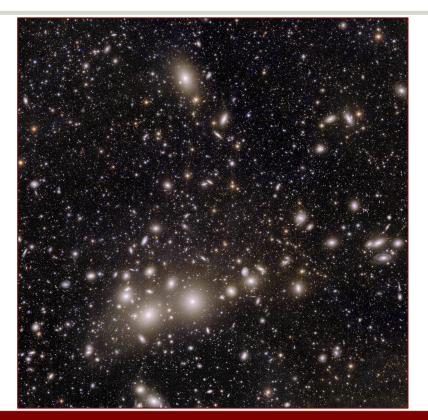
How Astrophysicists see the Universe

Christoph Welling



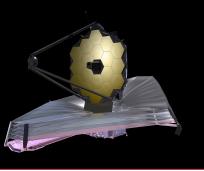
The Problem with Astrophysics



Telescopes!

 It is a great time to be an astrophysicist right now!

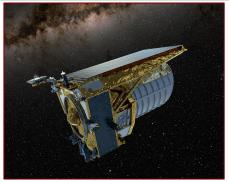
 Lots of new observatories operational or coming up soon



James Webb Space Telescope



Very Large Telescope



Euclid



Vera C. Rubin Observatory

So many pretty pictures!

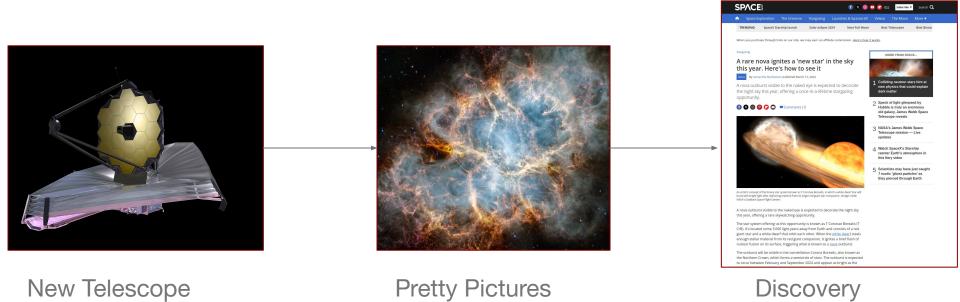


Perseus Cluster (Euclid)

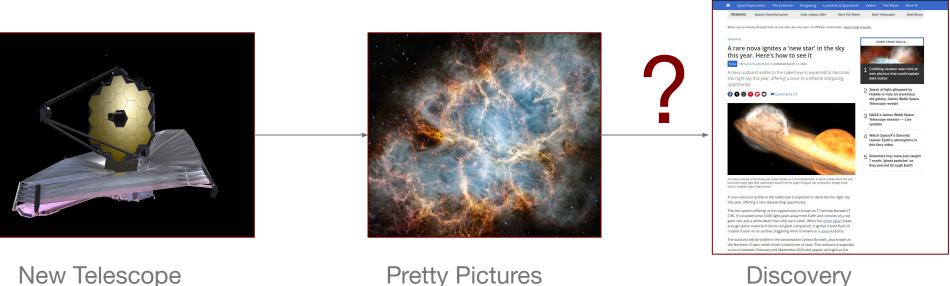


Crab Nebula (JWST)

The Oversimplified Story



The Oversimplified Story



SPACE

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New Telescope

Pretty Pictures

Not just about pretty pictures

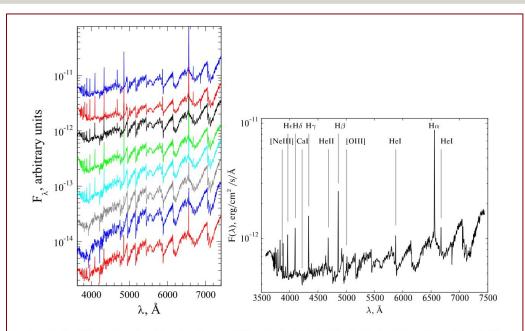


Figure 4: The observed spectra for T CrB (left panel) obtained on (from top to bottom) 22.03.2020, 05.01.2022, 27.01.2023, 26.02.2023, 09.03.2023, 24.06.2023, 26.05.2023, 14.07.2023. The fluxes are multiplied by an arbitrary constant for clarity. The right panel shows the spectrum of T CrB obtained on 05.01.2022 with the most prominent lines identified.

More than just telescopes



ALMA



Pierre Auger Observatory



LOFAR



IceCube



H.E.S.S.



LIGO

What is this series going to be about?

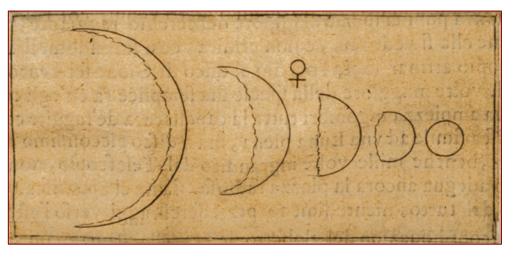
"How do we find out about things in space?"

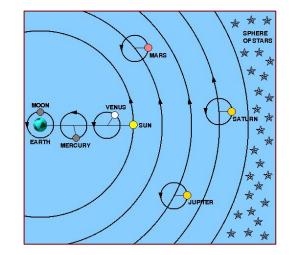
- 1. Introduction & Telescopes
- 2. Spectroscopy
- 3. Distances
- 4. Microwave & Radio Astronomy
- 5. Cosmic Rays
- 6. Gamma Rays
- 7. Neutrino Astronomy
- 8. Gravitational Waves

Galileo Galilei



Galileo Galilei





Galileo Galilei



Observationes Desuitary 2 S. Joris. manett. 12 0 ** 30. mone **0 * 2. x6n: Q** * 3. mon 0 * * 3. Ho. r. *0 * q. mont. *0 ** 6. mand **0 * 8. mane H. 13. # * * () Lo. mape. * * 0 * 11. * 0 # * 12. H. q regs: 0 * * 13. mane' *.0 * # 14 dane. * * * 0 *

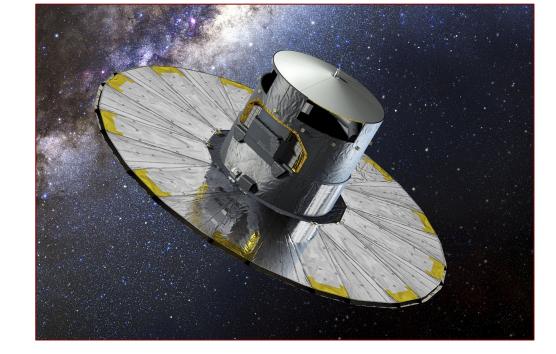
Tycho Brahe & Johannes Kepler



Tycho Brahe

Tycho Brahe & Johannes Kepler





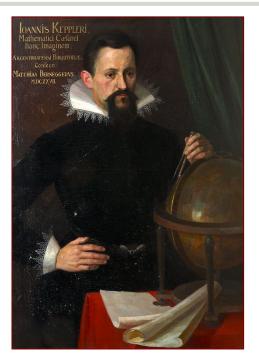
Tycho Brahe



Tycho Brahe & Johannes Kepler



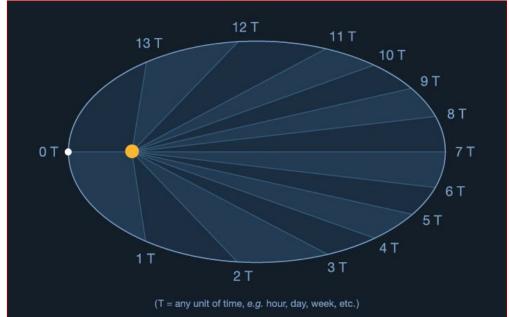
Tycho Brahe



Johannes Kepler

Kepler's Laws of Motion

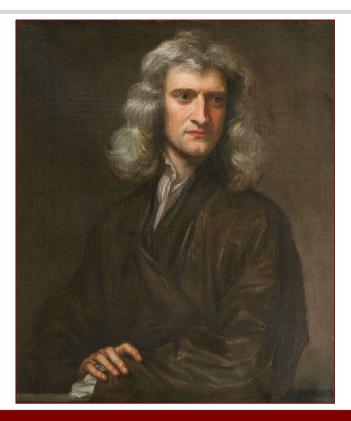
- 1. Orbits are ellipses, with the sun at one focus.
- 2. The line between sun and planet always sweeps the same area per time.
- 3. The square of the orbital period is proportional to the cube of the distance.



Isaac Newton

- Newton's Laws of Motion:
 - An object moves in a straight line, at constant speed, unless a force acts on it.
 - F = m * a
 - For every force, there is an equal, but opposite, reaction force.
- Newton's theory of Gravity:

$$F = G * m_1 * m_2 / d^2$$

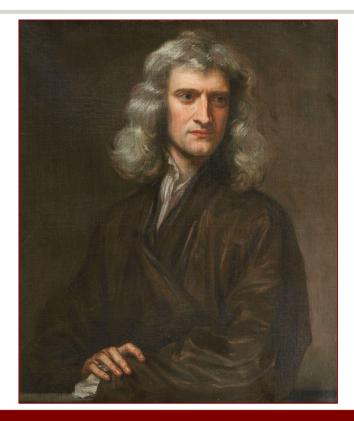


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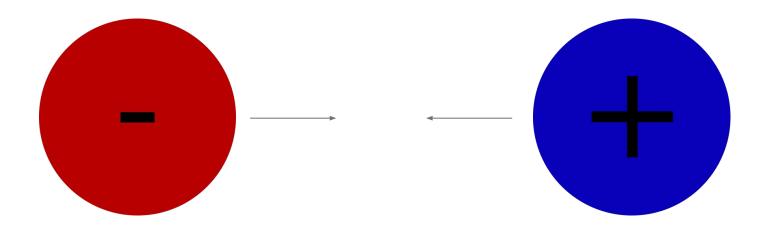
The planets are following the same laws as things on earth!







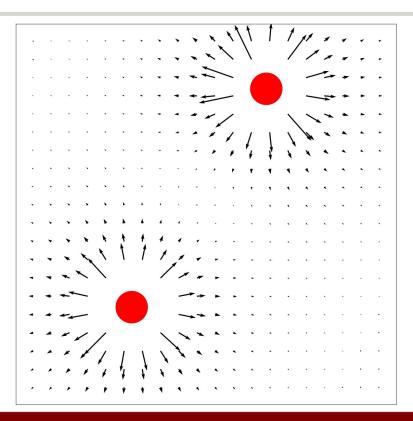
What is Light exactly?



What is Light exactly?



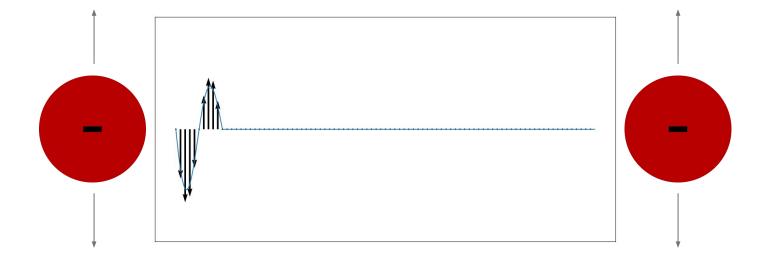
Electric Fields



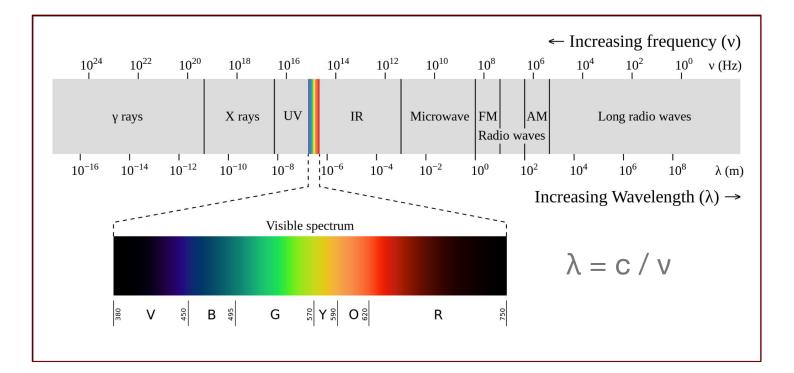
Electromagnetic Waves



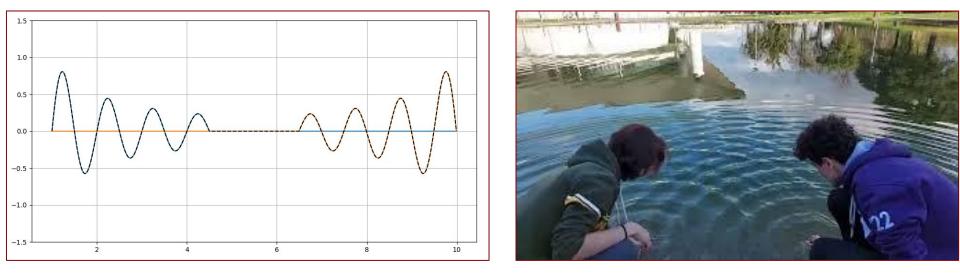
Electromagnetic Waves



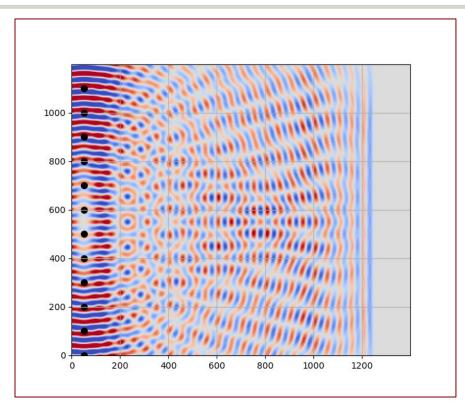
Electromagnetic Waves



Interference

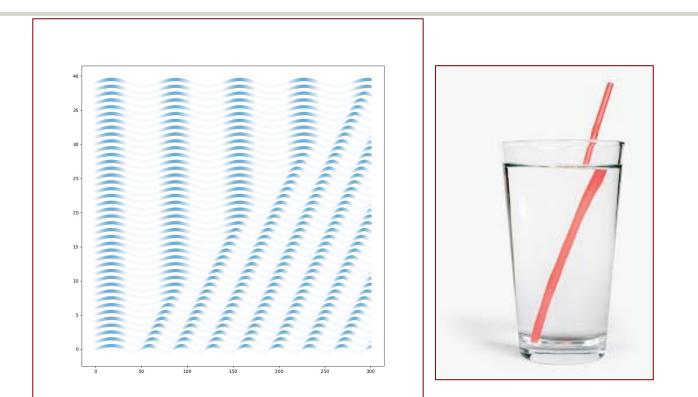


Wave optics

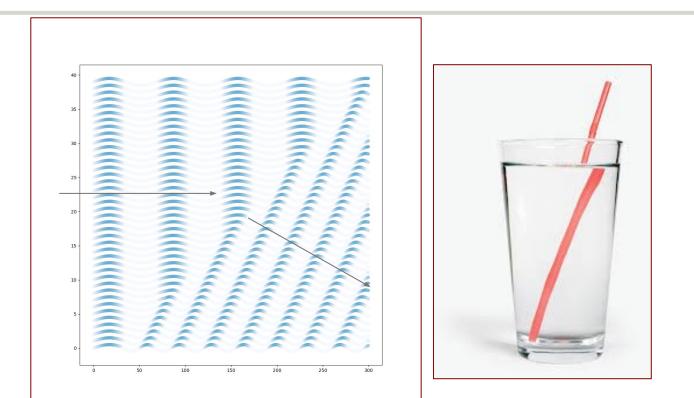


THE UNIVERSITY OF Arthur H. Compton Lecture Series

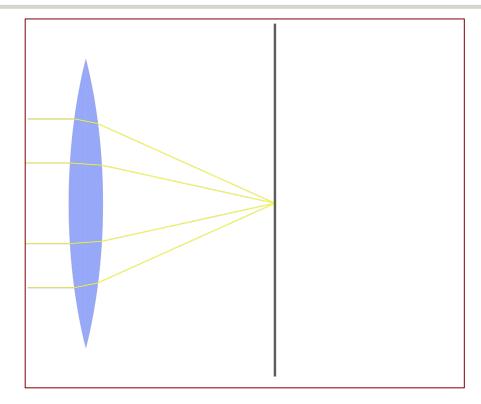
Refraction



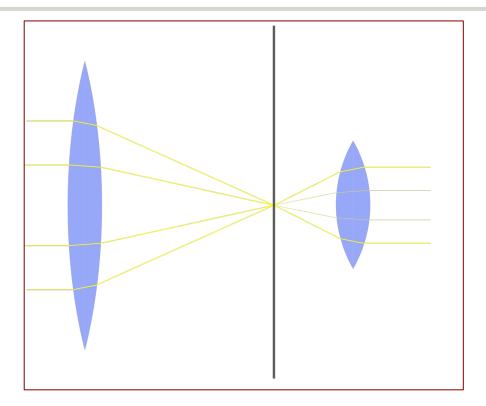
Refraction



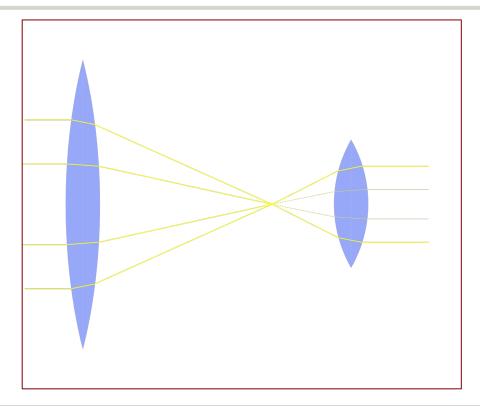
Ray optics and lenses



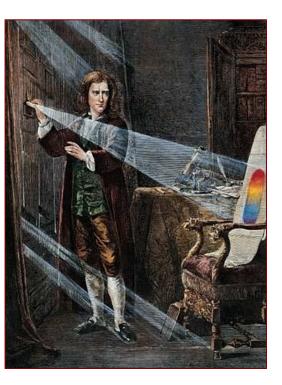
Ray optics and lenses

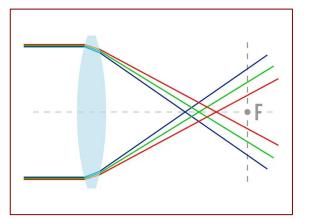


Ray optics and lenses



Chromatic Aberration



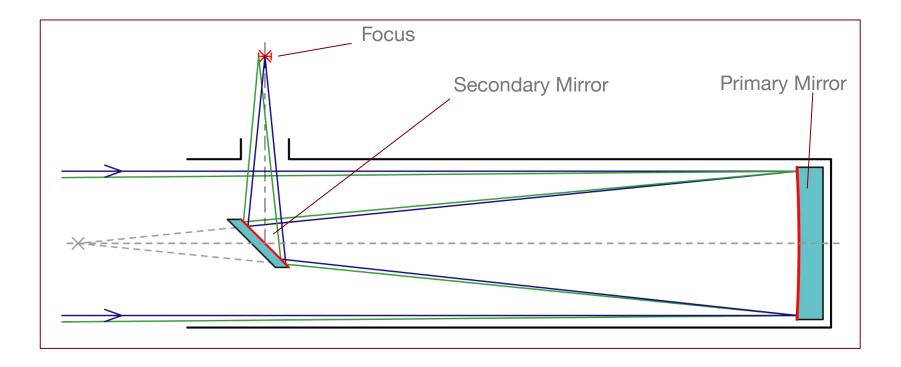




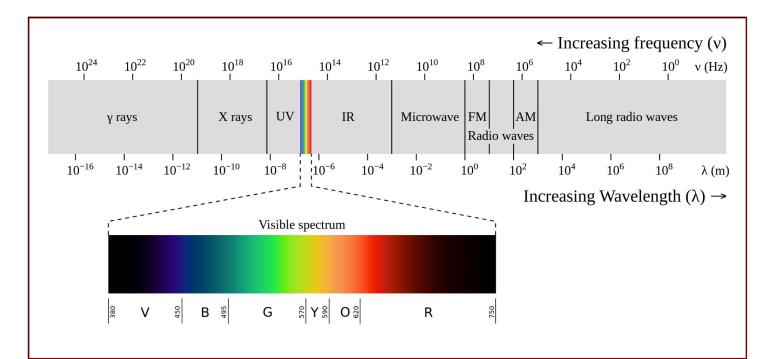
This text is hard to read!

THE UNIVERSITY OF Arthur H. Compton Lecture Series

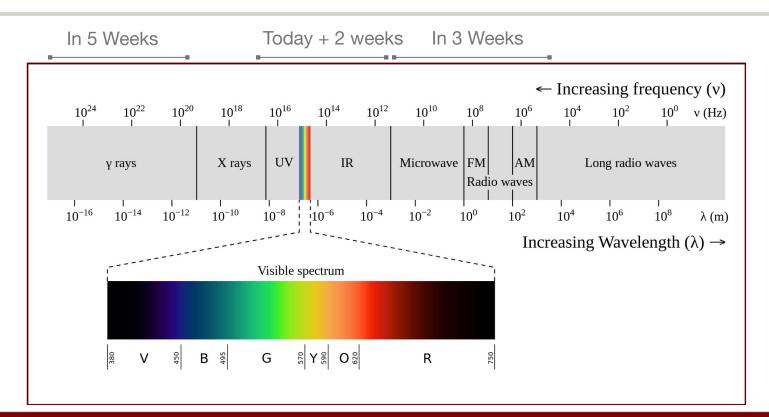
Reflector Telescopes



What Wavelengths do you want to see?

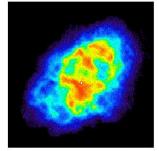


What Wavelengths do you want to see?



What Wavelengths do you want to see?

Crab Nebula: Remnant of an Exploded Star (Supernova)



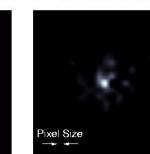
Radio wave (VLA)



Infrared radiation (Spitzer)



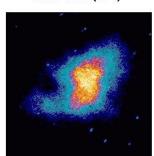
Visible light (Hubble)



High-energy X-ray (HEFT) *** 15 min exposure ***



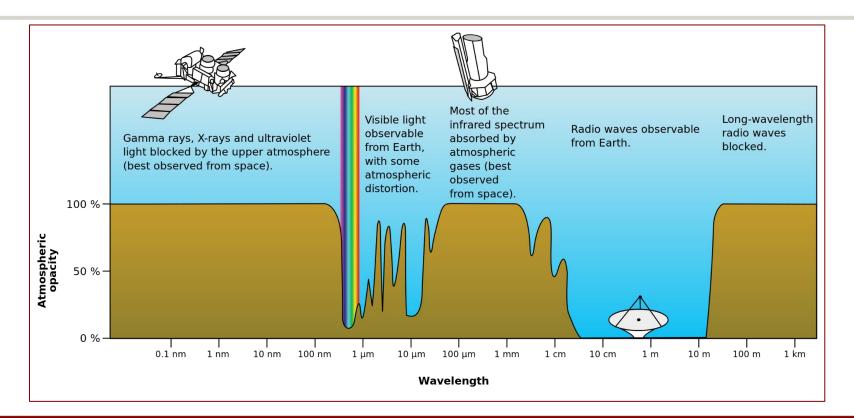
William Parsons



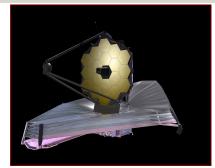
Ultraviolet radiation (Astro-1)

Low-energy X-ray (Chandra)

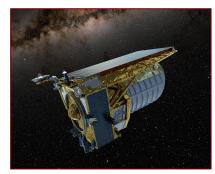
Atmospheric Absorption



What about space telescopes?



\$10 billion

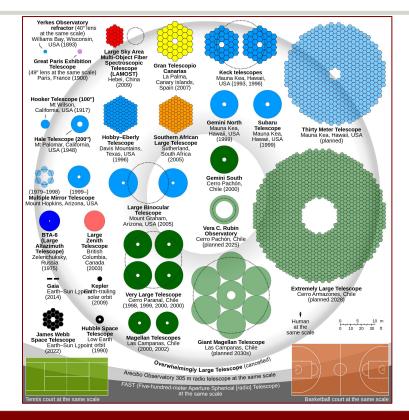




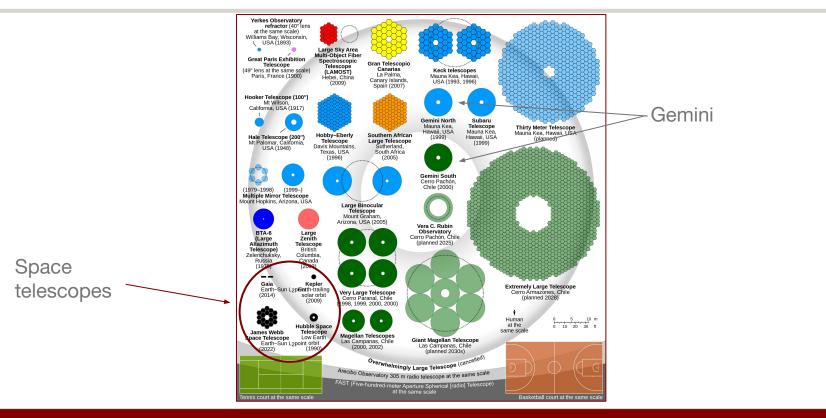


Gemini: \$187 million (2 telescopes)

Telescope Sizes

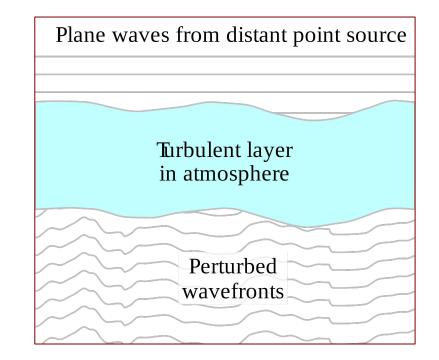


Telescope Sizes

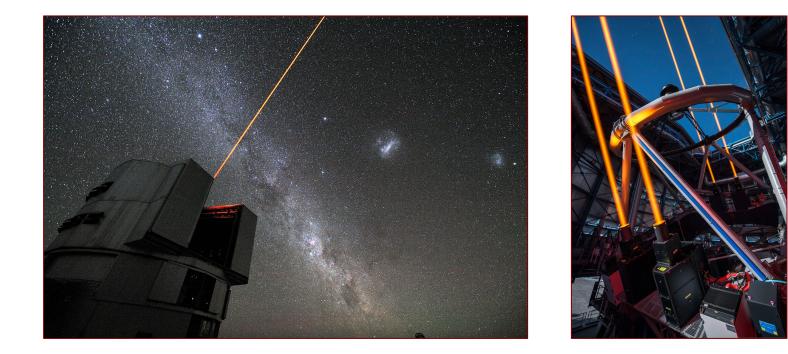


"Seeing"

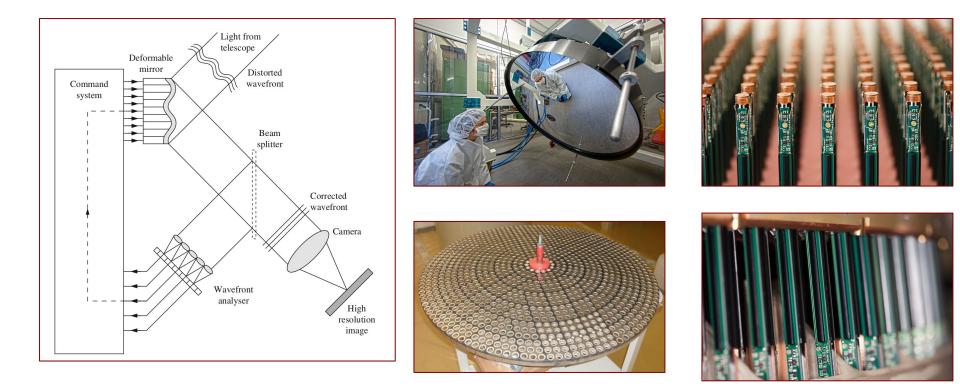




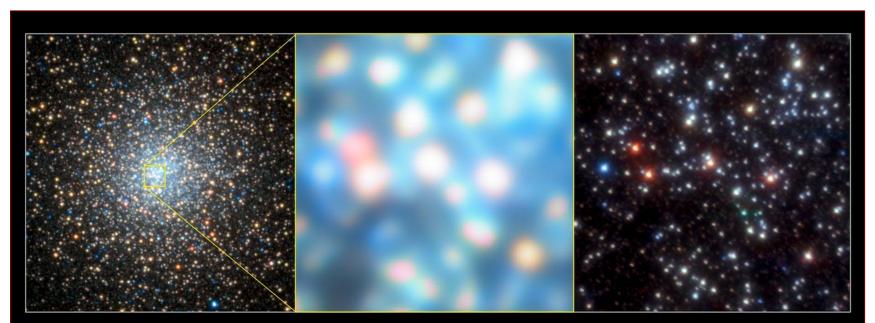
Adaptive Optics



Adaptive Optics



Adaptive Optics



VLT+MUSE Wide Field Mode without Adaptive Optics

VLT+MUSE Narrow Field Mode with Adaptive Optics

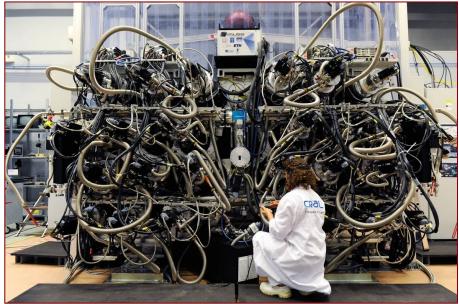
Instruments: VLT Unit Telescope 4

- High Acuity Wide-field K-band Imager
 - Near infrared
 - Wide field of view



Instruments: VLT Unit Telescope 4

- High Acuity Wide-field K-band Imager
 - Near infrared
 - Wide field of view
- Multi Unit Spectral Explorer
 - Visible to near infrared
 - Integral field spectroscopy

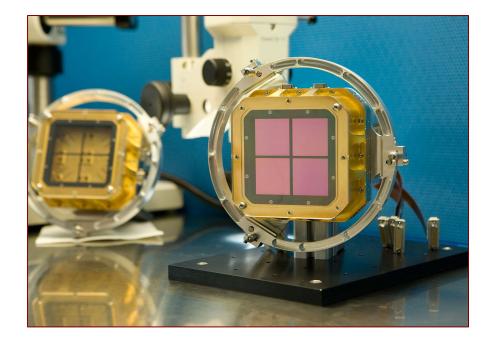


Instruments: VLT Unit Telescope 4

- High Acuity Wide-field K-band Imager
 - Near infrared
 - Wide field of view
- Multi Unit Spectral Explorer
 - Visible to near infrared
 - Integral field spectroscopy
- Enhanced Resolution Imager and Spectrograph
 - Near infrared
 - Integral field spectroscopy
 - Imaging
 - Coronagraph

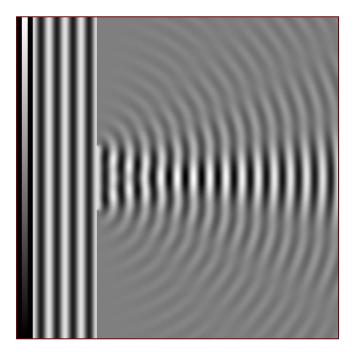


Charge Coupled Devices



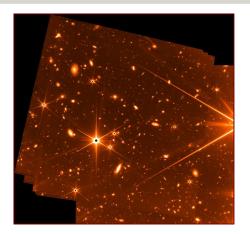


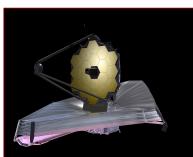
Diffraction



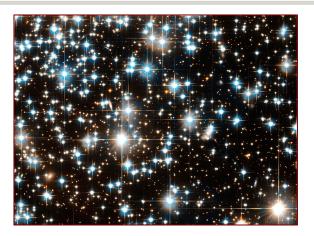


Diffraction



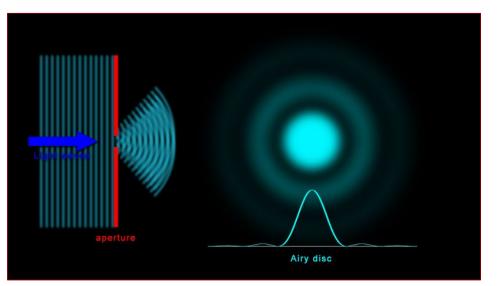




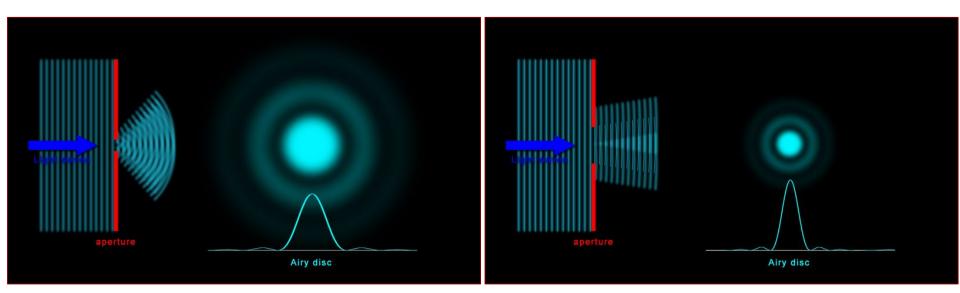






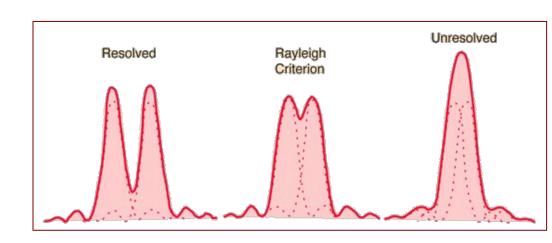






Resolution limit from diffraction:

 $\boldsymbol{\theta}\approx 1.22\;\lambda\,/\,D$

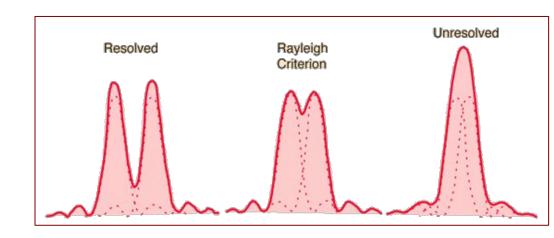


Resolution limit from diffraction:

 $\theta \approx 1.22 \ \lambda / D$

• Angular size of object:

 $\theta \approx r/d$



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 $\theta \approx r/d$

- Example: Nearby star with VLT
 - 8m mirror diameter
 - ~ 1µm wavelength
 - 10 lightyear distance

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Required size: 1.4*10¹⁰m

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Sun diameter: 1.4*10⁹m

Distance earth-sun: 1.5*10¹¹m

See you next week!

